THE FROISSART LIMIT....

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DO WE NEED FASTEST INCREASE OF THE SIGMA-TOT?

EXAMPLE: FIRST COLLISION IN COSMIC RAY SHOWER

$$P(X) = \frac{1}{\lambda} \exp\left(-\frac{X}{\lambda}\right)$$
 with $\lambda = \frac{1}{\sigma}$

$$\longrightarrow$$
 $\langle X \rangle = \lambda$

$$\longrightarrow SQRT(\langle X^2 \rangle - \langle X \rangle^2) = \lambda$$

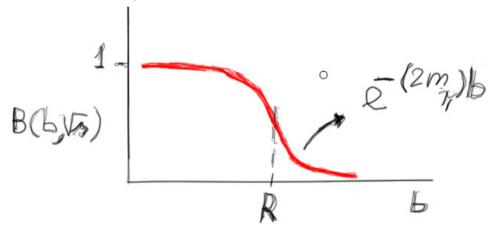
THIS IS IN THE DIRECTION OF AUGER RESULTS

ORIGIN OF INCREASE?

- •HEAVIER BEAM: p → Fe
- •INTRINSIC INCREASE OF σ_{pp}

THE HEISENBERG-FROISSART ARGUMENT (50-60'S)

IN QCD THE LARGE DISTANCE LIMIT IS NOT CONTROLLED BY THE GLUONS (1/R POTENTIAL) BUT BY THE LIGHTEST EXTENDED HADRON (YUKAVA POTENTIAL, ROLE OF THE PION)



$$\sqrt{s} \exp(-2m_{\pi}R) \sim \sqrt{s_0} \geq m_{\pi}$$

$$R < \frac{1}{4m_{\pi}} \ln \left(\frac{s}{s_0} \right)$$

$$\sigma = \pi R^2 \longrightarrow$$

$$\sigma < \frac{\pi}{(4m_{\pi})^2} \ln^2 \left(\frac{s}{s_0}\right)$$

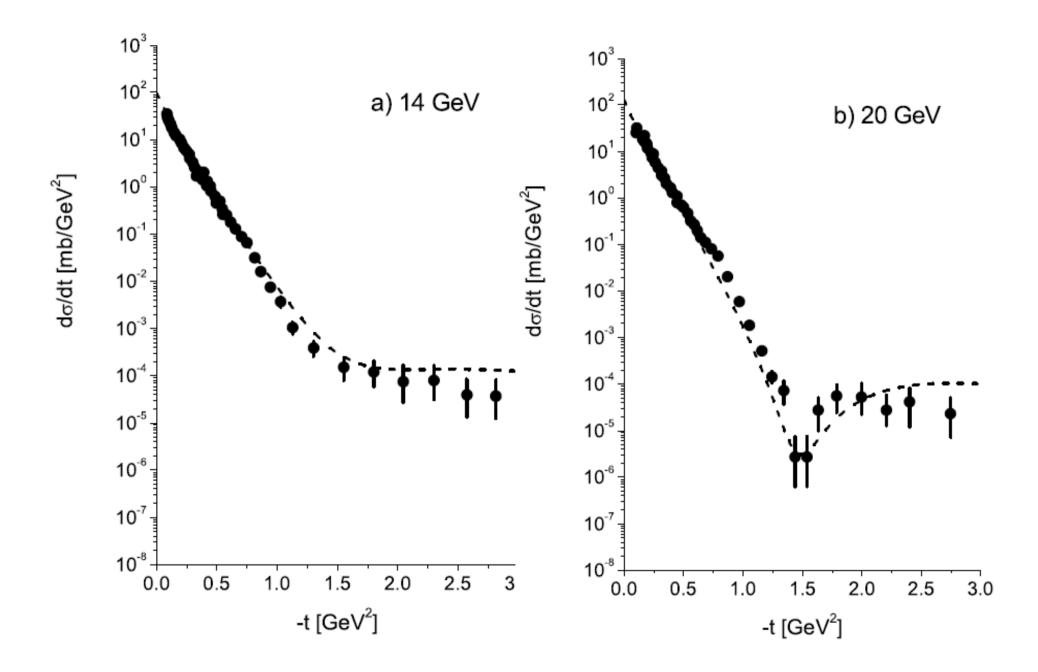
NON LINEAR MODEL FOR THE HIGH ENERGY AMPLITUDE

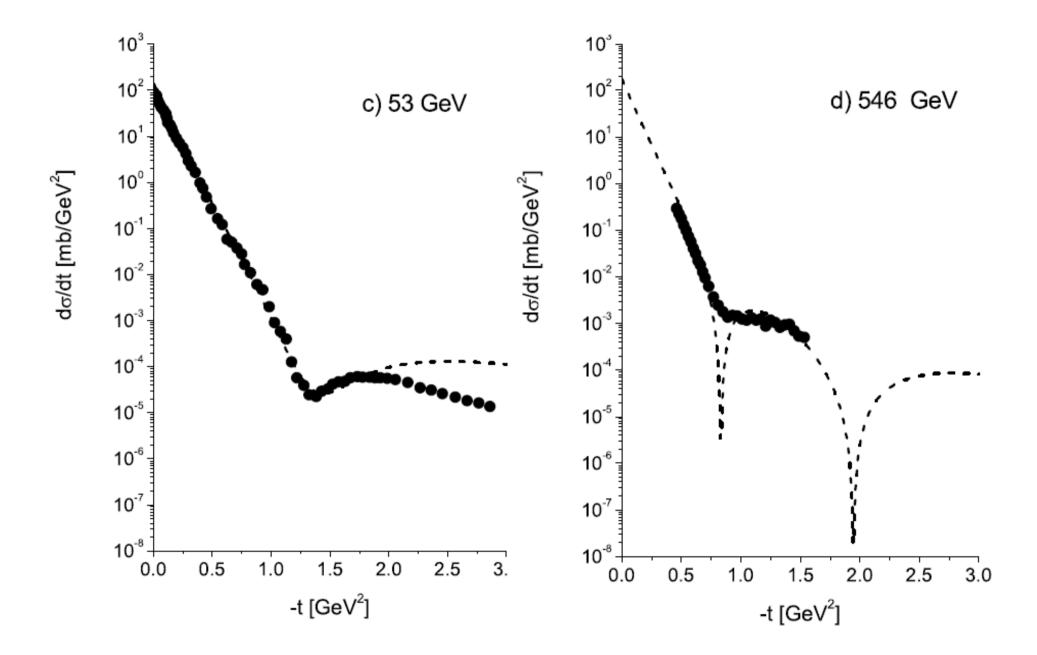
LOGISTIC EQUATION WITH SATURATION (OF UNITARITY → BLACK DISK)

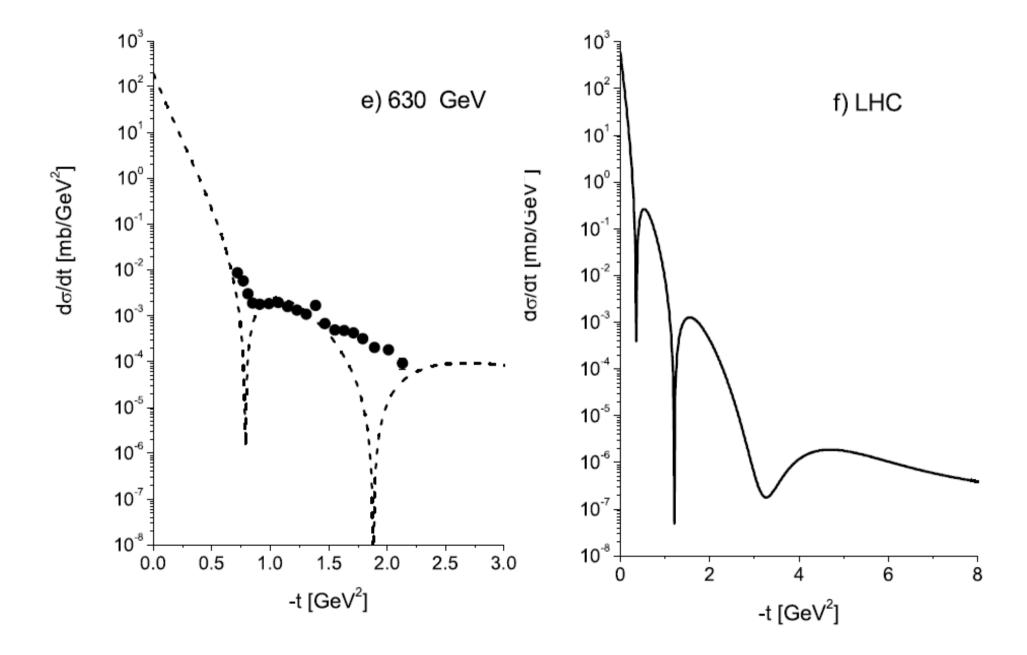
B(b,R) R: ENERGY DEPENDENT PARAMETER

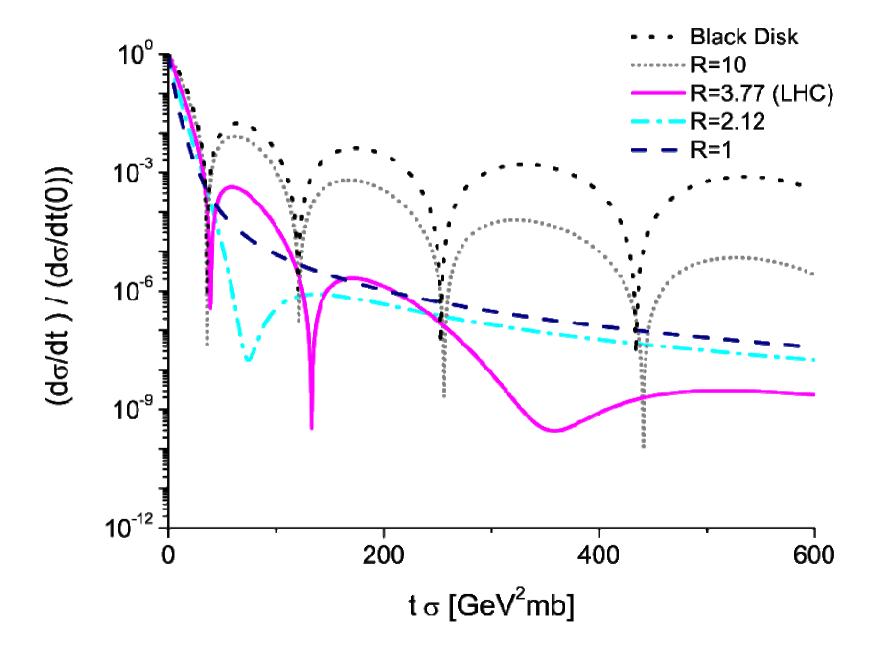
[FROM $\exp\left(-\frac{b}{\gamma}\right)$ (AT SMALL R) TO BLACK DISK (AT LARGE R)]

RESULTS:





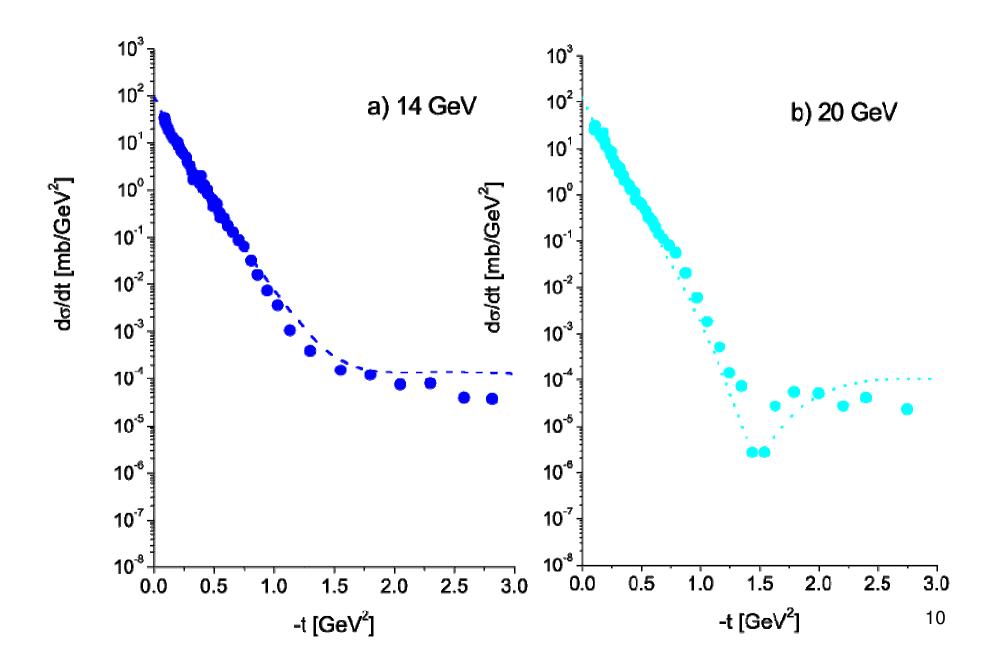


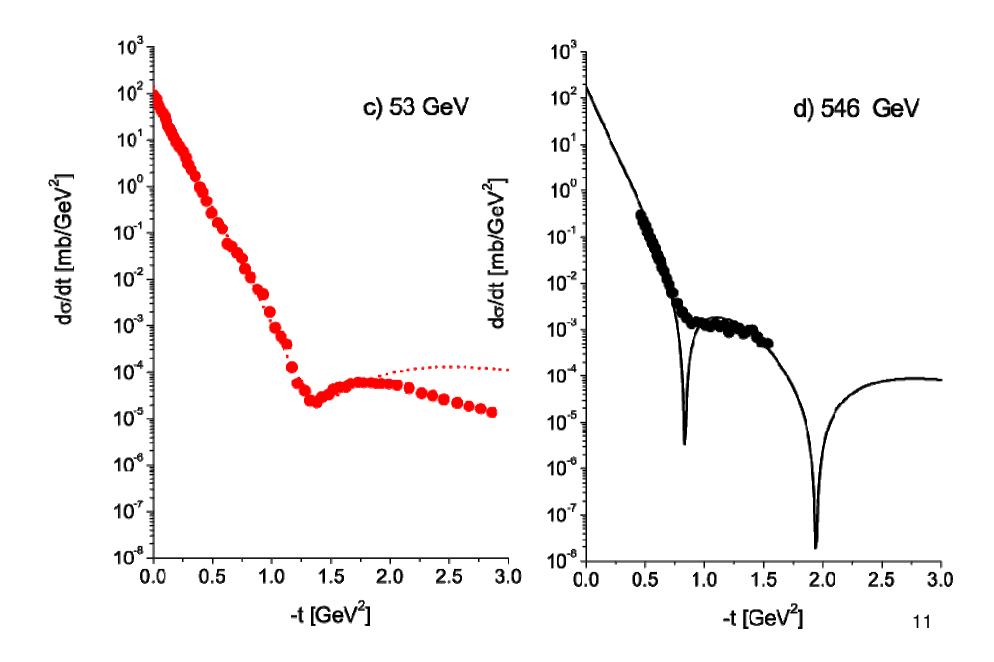


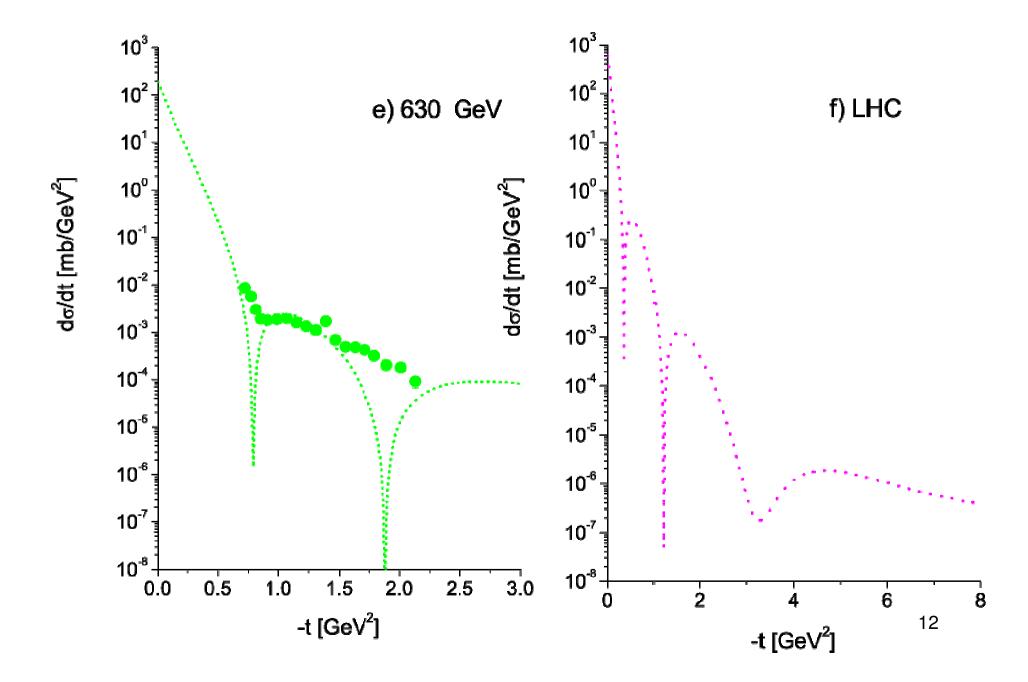
CONCLUSION: IS FROISSART BOUND VIOLATED?

THERE IS NO INDICATION

- CROSS-SECTION MAY INCREASE WITHOUT VIOLATION OF THE FROISSART BOUND [ULRICH, JAIME ALVAREZ-MUNIZ]
- → NUSSINOV IDEA
 GLUE BALLS BEFORE PIONS...







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, with $\lambda=1/\sigma$

$$\rightarrow$$
 $\langle X \rangle = \lambda$

$$\longrightarrow$$
 SQRT($<$ X²>- $<$ X>²)= λ

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