

# **II Workshop on Black Holes**

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# Zero frequency limit of gravitational radiation in high-energy collisions

Madalena Lemos (CENTRA/IST)

Work in progress:

E. Berti, V. Cardoso, T. Hinderer, M.L., F. Pretorius, U. Sperhake, N. Yunes

Zero frequency limit of gravitational radiation in high-energy collisions

# **Zero Frequency Limit (ZFL)** (Weinberg '64 Smarr '77)

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

$$h_{\mu\nu} = 4G \int \frac{S_{\mu\nu} (t_{ret}, \mathbf{x}')}{|\mathbf{x}' - \mathbf{x}|} d^3 \mathbf{x}' \qquad S_{\mu\nu} = T_{\mu\nu} - \frac{1}{2} \eta_{\mu\nu} T$$

L. Smarr, Phys. Rev. D 15, 2069 (1977)

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- Hard collision → incoming and outgoing trajectories with constant velocities

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# Zero Frequency Limit (ZFL)

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- Zero duration → infinite acceleration
- Hard collision → incoming and outgoing trajectories with constant velocities
  - $\rightarrow$  Valid for arbitrary velocities
  - $\rightarrow$  Details of internal structure irrelevant
  - $\rightarrow$  Flat, large memory

$$\frac{dE}{d\omega} \propto \left(h_{t=\infty} - h_{t=-\infty}\right)^2$$

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 $\rightarrow$  Free particles, changing abruptly at t=0:



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$$T^{\mu\nu}(\mathbf{x},t) = \sum_{i=1}^{2} \frac{p_{i}^{\mu} p_{i}^{\nu}}{E_{i}} \delta^{3}(\mathbf{x} - \mathbf{v}_{i}t)\theta(-t) + \sum_{i=1}^{2} \frac{p_{i}^{\prime\mu} p_{i}^{\prime\nu}}{E_{i}^{\prime}} \delta^{3}(\mathbf{x} - \mathbf{v}_{i}^{\prime}t)\theta(t)$$

$$\stackrel{\mathbf{z}^{\prime}}{\stackrel{\mathbf{z}^{\cdot}}{\stackrel{\mathbf{z}^{\cdot}}{\stackrel{\mathbf{z}^{\cdot}}{\stackrel{\mathbf{z}^{\cdot}}{\stackrel{\mathbf{z}^{\cdot}}{\stackrel{\mathbf{z}^{\cdot}}{\stackrel{\mathbf{z}^{\cdot}}{\stackrel{\mathbf{z}^{\cdot}}{\stackrel{\mathbf{z}^{\cdot}}{$$

- $\rightarrow$  Independent of  $\omega$
- $\rightarrow$  Introduce a **cutoff** frequency

L. Smarr, Phys. Rev. D 15, 2069 (1977)

 $\frac{dP^{i}}{d\omega} = \int_{\Omega} \frac{d^{2}E}{d\omega d\Omega} n^{i} d\Omega$ **Radiated Momentum:**  $\frac{dP^i}{d\omega} = 0 \quad , i = y, z \longrightarrow$ Momentum is radiated in the direction of motion  $\frac{dP^x}{d\omega} = \frac{m_2^2}{\pi v_2^2 (v_1 + v_2)} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{2} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{2} \right) + \frac{dP^x}{2} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{2} \right)$ +  $\frac{m_1^2}{\pi v_1^2 (v_1 + v_2)} (v_1^2 (v_1 - v_2) - 5v_1 - 3v_2) \operatorname{arctanh}(v_1) +$ +  $\frac{\gamma_1^2 m_1^2}{\pi v_1 v_2^3} (v_1 - v_2) \left( v_1^2 \left( v_2^2 - 3 \right) - 5v_1 v_2 - 3v_2^2 \right)$ 

 $\frac{dP^{i}}{d\omega} = \int_{-\infty}^{\infty} \frac{d^{2}E}{d\omega d\Omega} n^{i} d\Omega$ **Radiated Momentum:**  $\frac{dP^i}{d\omega} = 0 \quad , i = y, z \longrightarrow$ Momentum is radiated in the direction of motion  $\frac{dP^x}{d\omega} = \frac{m_2^2}{\pi v_2^2 (v_1 + v_2)} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{\pi v_2^2 (v_1 + v_2)} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{\pi v_2^2 (v_1 + v_2)} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{\pi v_2^2 (v_1 + v_2)} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{\pi v_2^2 (v_1 + v_2)} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{\pi v_2^2 (v_1 + v_2)} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{\pi v_2^2 (v_1 + v_2)} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{\pi v_2^2 (v_1 + v_2)} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{\pi v_2^2 (v_1 + v_2)} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{\pi v_2^2 (v_1 + v_2)} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{\pi v_2^2 (v_1 + v_2)} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{\pi v_2^2 (v_1 + v_2)} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{\pi v_2^2 (v_1 + v_2)} \left( \left( v_1 \left( v_2^2 + 3 \right) \right) - v_2^3 + 5v_2 \right) \operatorname{arctanh}(v_2) + \frac{dP^x}{\pi v_2^2 (v_1 + v_2)} \right) \right)$ +  $\frac{m_1^2}{\pi v_1^2(v_1+v_2)}(v_1^2(v_1-v_2)-5v_1-3v_2)\operatorname{arctanh}(v_1) +$ +  $\frac{\gamma_1^2 m_1^2}{\pi v_1 v_3^3} (v_1 - v_2) \left( v_1^2 \left( v_2^2 - 3 \right) - 5v_1 v_2 - 3v_2^2 \right)$  $= \begin{cases} 0, & m_1 = m_1 \\ \frac{m_1^2 \gamma_1^2 \left( v_1 \left( 15 - 13v_1^2 \right) - 3 \left( v_1^4 - 6v_1^2 + 5 \right) \operatorname{arctanh}(v_1) \right)}{3\pi v^2}, & m_1 \ll m_2 \end{cases}$ 



Point-particle falling into a black hole



Point-particle falling into a black hole

Collision of two equal mass black holes

U. Sperhake, V. Cardoso, F. Pretorius, E. Berti, and

#### **Center of momentum frame**



#### **Center of momentum frame**



 $\nabla_{\mu}T^{\mu\nu} \neq 0 \text{ for } \nu = x, y \longrightarrow$  Energy momentum tensor is not conserved!

#### Constraining forces:

(Price and Sandberg '73)

$$T_{\text{tens}}^{xx}(t, \mathbf{x}) = -\tau_1(r)S^2\delta(\cos\theta)\delta(\phi + \Omega t - \pi/2)\Theta(t) - \tau_2(r)S^2\delta(\cos\theta)\delta(\phi + \Omega t - 3\pi/2)\Theta(t)$$

$$T_{\text{tens}}^{yy}(t, \mathbf{x}) = -\tau_1(r)C^2\delta(\cos\theta)\delta(\phi + \Omega t - \pi/2)\Theta(t) - \tau_2(r)C^2\delta(\cos\theta)\delta(\phi + \Omega t - 3\pi/2)\Theta(t)$$

$$T_{\text{tens}}^{xy}(t, \mathbf{x}) = -\tau_1(r) S C \delta(\cos \theta) \delta(\phi + \Omega t - \pi/2) \Theta(t) - \tau_2(r) S C \delta(\cos \theta) \delta(\phi + \Omega t - 3\pi/2) \Theta(t)$$

$$\tau_1(r) = \frac{m_1 \xi \Omega^2}{r^2 \sqrt{1 - (\xi \Omega)^2}}, \quad \tau_2(r) = \frac{m_2 \xi_2 \Omega^2}{r^2 \sqrt{1 - (\xi_2 \Omega)^2}}$$

R. H. Price and V. D. Sandberg, Phys. Rev. D 8, 1640 (1973)

**Radiated energy:** 
$$\frac{d^2 E}{d\omega d\Omega} = 2G\omega^2 \left( T^{\mu\nu}(\mathbf{k},\omega) T^*_{\mu\nu}(\mathbf{k},\omega) - \frac{1}{2} \left| T^{\lambda}_{\ \lambda}(\mathbf{k},\omega) \right|^2 \right)$$

Fourier transform:

$$\begin{split} T^{\mu\nu}(\mathbf{k},\omega) &= \frac{p_{1}^{\mu}p_{1}^{\nu}}{2\pi i E_{1}(\omega-v_{1}k_{x})}e^{-ik_{y}\frac{b}{2}} + \frac{p_{2}^{\mu}p_{2}^{\nu}}{2\pi i E_{2}(\omega+v_{2}k_{x})}e^{ik_{y}\frac{b}{2}} + \\ &+ \sum_{i=1}^{2}\int_{-\infty}^{\infty}\frac{p_{i}^{\prime\mu}(t)p_{i}^{\prime\nu}(t)}{2\pi E_{i}^{\prime}(t)}\exp(i\omega t - i\mathbf{k}\cdot\mathbf{x}_{i}^{\prime}(t))\Theta(t)dt + \\ &+ \frac{1}{2\pi}\int_{-\infty}^{\infty}dt\int_{-\infty}^{\infty}dx\int_{-\infty}^{\infty}dy\int_{-\infty}^{\infty}dz\,T_{\text{tens}}^{\mu\nu}(\mathbf{x},t)e^{i\omega t - i\mathbf{k}\cdot\mathbf{x}} \end{split}$$

**Radiated energy:** 
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$$e^{i\omega t - i\mathbf{k}\cdot\mathbf{x}_{i}'(t)} = e^{i\omega t} \exp\left(\pm i\frac{\omega b}{2}\sin\theta\sin(\Omega t + \phi)\right)$$
$$e^{iz\sin(\alpha)} = \sum_{n=-\infty}^{n=+\infty} J_{n}(z)e^{in\alpha} \quad ---- \text{Jacobi-Anger expansion}$$

 $N \geq 10$  sufficient for an accuracy of 1% or better



$$m_1 = m_2 \Longrightarrow \Omega = \frac{2v_1}{b}$$

• All spectra blow up for  $\omega = 2n\Omega$ , n = 1, 2, ...Except for  $\hat{\mathbf{k}} = \mathbf{e}_z$   $\rightarrow$  Only blows up at  $\omega = 2\Omega$ *E. Poisson, Phys. Rev. D47, 1497 (1993).* 

#### Zero frequency limit:

 $\rightarrow$  Smarr's head-on collision is recovered

$$\frac{d^2 E}{d\omega d\Omega} = \frac{Gm_1^2 \gamma_1^2 v_1^4 \left(\sin^2 \theta \cos^2 \phi - 1\right)^2}{\pi^2 \left(v_1^2 \sin^2 \theta \cos^2 \phi - 1\right)^2}$$



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Zero frequency limit:

$$\left. \frac{d^2 E}{d\omega d\Omega} \right|_{\hat{\mathbf{k}} \neq \mathbf{e}_{\mathbf{x}}} \sim (b\omega)^2$$

$$\left. \frac{d^2 E}{d\omega d\Omega} \right|_{\mathbf{\hat{k}} = \mathbf{e}_{\mathbf{x}}} \sim (b\omega)^4$$

#### ZFL: Non Head-on Collision – extreme mass ratio



#### Center of momentum frame

$$\mu \equiv m_1 \ll m_2 \equiv M$$

• All spectra blow up for

$$\omega = n\Omega$$
,  $n = 1, 2, ...$   
Except for  $\hat{\mathbf{k}} = \mathbf{e}_z$   
 $\rightarrow$  Only blows up at  $\omega = 2\Omega$ 

$$\Omega = \frac{\gamma_1 \mu v_1 \xi + \gamma_2 M v_2 (b - \xi)}{\gamma_1 \mu \xi + \gamma_2 M (b - \xi)}$$
$$\xi = \frac{b \gamma_2 M}{\gamma_1 \mu + \gamma_2 M}$$

#### ZFL: Non Head-on Collision – extreme mass ratio



Zero frequency limit:

 $\rightarrow$  Smarr's head-on collision is recovered

$$\frac{d^2 E}{d\omega d\Omega} = \frac{G\mu^2 \gamma_1^2 v_1^4 \left(\sin^2\theta \cos^2\phi - 1\right)^2}{4\pi^2 \left(1 - v_1 \sin\theta \cos\phi\right)^2}$$

#### ZFL: Non Head-on Collision – extreme mass ratio



Zero frequency limit:

 $\rightarrow$  Smarr's head-on collision is recovered

$$\left. \frac{d^2 E}{d\omega d\Omega} \right|_{\hat{\mathbf{k}} \neq \mathbf{e}_{\mathbf{x}}} \sim (b\omega)^2 \quad \left. \frac{d^2 E}{d\omega d\Omega} \right|_{\hat{\mathbf{k}} = \mathbf{e}_{\mathbf{x}}} \sim (b\omega)^4$$



- $\rightarrow$  Head-on limit spectrum is flat for small frequencies
- → Different behaviour near resonance for numerical data and toy model
- $\rightarrow M\omega \rightarrow 0$  limit very weakly dependent on  $L_Z/L_{crit}$  consistent with the toy model



 $\rightarrow M\omega \rightarrow 0$  limit very weakly dependent on  $L_Z/L_{crit}$ 

$$ightarrow$$
 Head-on case – ZFL yields  $\lim_{\omega \to 0} \frac{dE_{22}/d\omega}{dE_{20}/d\omega} = \frac{3}{2}$ 

 $\rightarrow$  Good agreement with numerical results



 $\rightarrow m = \pm l$  have positive/negative slope for  $M\omega \rightarrow 0$ 

$$\rightarrow \left| \frac{dE}{d\omega d\Omega} \right| \sim b$$
  
$$\rightarrow \text{Toy model:} \quad \left| \frac{dE}{d\omega d\Omega} \right| \sim b^2$$

#### Summary

- Zero-frequency limit is independent of the impact parameter
- Numerical results weakly dependent on the impact parameter
- Different small frequency behaviour
- Not the same behaviour near resonances

#### **Future Work**

- $\rightarrow$  Dependence on the constraining forces
- $\rightarrow$  Particles with structure
- $\rightarrow$  Higher-dimensions
- → Radiated Momentum
- $\rightarrow$  Different ways to model the collision

#### Thank you

# **Backup Slides**



Small M<sub>w</sub>:

$$\frac{d^2 E}{d\omega d\Omega}\Big|_{\hat{\mathbf{k}}\neq\mathbf{e}_{\mathbf{x}}}\sim(b\omega)^2\qquad\qquad\qquad \frac{d^2 E}{d\omega d\Omega}\Big|_{\hat{\mathbf{k}}=\mathbf{e}_{\mathbf{x}}}\sim(b\omega)^4$$





Same structure

Zero frequency limit:

 $\rightarrow$  Smarr's head-on collision is recovered

$$\frac{d^2 E}{d\omega d\Omega} = \frac{Gm^2 \gamma_1^2 v_1^2 \left(\sin^2(\theta) \cos^2(\phi) - 1\right)^2 (v_1 + v_2)^2}{4\pi^2 \left(1 - v_1 \sin(\theta) \cos(\phi)\right)^2 \left(1 + v_2 \sin(\theta) \cos(\phi)\right)^2}$$



 $\rightarrow m = \pm l$  have positive/negative slope for  $M\omega \rightarrow 0$ 

$$\rightarrow \left| \frac{dE}{d\omega d\Omega} \right| \sim b$$
  
$$\rightarrow \text{Toy model:} \quad \left| \frac{dE}{d\omega d\Omega} \right| \sim b^2$$



 $\rightarrow$  Head-on case – ZFL yields  $\lim_{\omega \to 0} \frac{dE_{22}/d\omega}{dE_{20}/d\omega} = \frac{3}{2}$ 

 $\rightarrow$  Good agreement with numerical results

## Smarr

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu} \qquad \cos(\theta') = \sin(\theta)\cos(\phi)$$



$$\frac{dE}{d\omega d\Omega} = \frac{M^2 \gamma^2 v^4}{\pi^2} \frac{\sin^4 \theta'}{(1 - v^2 \cos^2 \theta')^2}$$