## Acceleration in Special Relativity and Mach's Principle

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## Equivalence principle

Weak Equivalence Principle: all particles fall at the same rate in a gravitational field, independent of their mass and composition


The equivalence principle: In a freely falling (non-rotating) laboratory occupying a small region of spacetime, the laws of physics are those of special relativity.

## Special Relativity

The equivalence principle: In a freely falling (non-rotating) laboratory occupying a small region of spacetime, the laws of physics are those of special relativity.

Special Relativity = Flat Lorentzian Geometry!
Objects moving in the absence of gravitational field

General Relativity = Curved Lorentzian Geometry!
Objects moving in the presence of gravitational field


## Flat space (Riemannian) = Euclidean geometry



Pato Donald no país da MateMágica!

$$
d z^{2}=d x^{2}+d y^{2}
$$

## Flat space (Lorentzian) = Minkowski space-time

Metric $c^{2} d^{2}=-c^{2} d^{2}+d x^{2}$

- Timelike vectors
- Spacelike vectors
- Null or Lightlike vectors



## Galilean view vs Einstein view of Past and Future




Image: General Relativity: An introduction for physicists. Hobson, Efstathiou \& Lasenby

## Simultaneity in Special Relativity



One can "read" the time dilation effect from the diagram without doing any computation


## Curves of constant acceleration

## Newtonian




$$
\frac{d x}{d t}=a t
$$

## Curves of constant acceleration

## Definitions

$$
\begin{aligned}
& \boldsymbol{x}=(t(\lambda), x(\lambda)) \\
& \dot{\boldsymbol{x}}=(d t / d \lambda, d x / d \lambda)
\end{aligned}
$$

$$
c d \tau=\sqrt{-\|\dot{\boldsymbol{x}}\|_{M i n k}^{2}}
$$

$$
\boldsymbol{u}=\frac{1}{c}\left(\frac{d t}{d \tau}, \frac{d x}{d \tau}\right)
$$

$$
\boldsymbol{a}=\frac{1}{c^{2}}\left(\frac{d^{2} t}{d \tau^{2}}, \frac{d^{2} x}{d \tau^{2}}\right)
$$

Spacetime diagram


Hyperbola in spt

$$
\begin{aligned}
x(\lambda) & =\frac{c^{2}}{a_{p}}\left(\cosh \left[\frac{a_{p}}{c} \lambda\right]-1\right) \\
t(\lambda) & =\frac{c}{a_{p}} \sinh \left[\frac{a_{p}}{c} \lambda\right]
\end{aligned}
$$

Proper time

Proper 4-velocity

Proper 4-acceleration

Worldline

Tangent vector

## Curves of constant acceleration

## Special Relativity



Hyperbola in spt diagram

$$
\begin{aligned}
& \dot{\boldsymbol{x}}=\left(\cosh \left[\frac{a_{p}}{c} \lambda\right], c \sinh \left[\frac{a_{p}}{c} \lambda\right]\right) \\
& \|\dot{\boldsymbol{x}}\|_{M i n k}^{2}=-c^{2} \Longrightarrow \tau=\lambda \\
& \frac{d x / d \lambda}{d t / d \lambda}=\frac{d x}{d t}=c \tanh \left[\frac{a_{p}}{c} \lambda\right] \\
& \boldsymbol{u}=\left(\frac{1}{c} \cosh \left[\frac{a_{p}}{c} \tau\right], \sinh \left[\frac{a_{p}}{c} \tau\right]\right) \\
& \boldsymbol{a}=\frac{1}{c}\left(\frac{a_{p}}{c^{2}} \sinh \left[\frac{a_{p}}{c} \lambda\right], \frac{a_{p}}{c} \cosh \left[\frac{a_{p}}{c} \lambda\right]\right)
\end{aligned}
$$



Hyperbolic worldline in spacetime
is that of constant proper 4-acceleration!

$$
\|\boldsymbol{a}\|_{\text {Mink }}^{2}=\frac{a_{p}^{2}}{c^{4}},
$$

## Twin Paradox



Tom and Bob are 21 years old

Tom travels at a speed
0.96 c to a distant star. Instantaneously upon arrival, returns to earth at speed 0.96 c.

Tom is 35 years old

Bob is 71 years old!

## Twin Paradox

$$
x(t)= \begin{cases}v t & \text { if } t \in[0, T / 2] \\ v(T-t) & \text { if } t \in[T / 2, T]\end{cases}
$$

$x(t)= \begin{cases}v t & \text { if } t \in[0, T / 2] \\ v(T-t) & \text { if } t \in[T / 2, T]\end{cases}$
Tom's worldline


$$
\|\dot{\boldsymbol{x}}\|_{\text {Mink }}^{2}=-c^{2} d t+( \pm v d t)^{2} \Longrightarrow c d \tau=\sqrt{c^{2}-v^{2}} d t
$$

$$
\begin{aligned}
& \frac{\Delta \tau}{\Delta t}=\sqrt{1-\frac{v^{2}}{c^{2}}} \\
& \frac{T^{\prime}}{T}=\sqrt{1-\frac{v^{2}}{c^{2}}} \leq 1
\end{aligned}
$$

## Triangle inequality (Euclidean space)



## Reversed triangle inequality (Minkowski space)



## Twin paradox as the reversed triangle inequality



$c>a+b$



David Bowie as Major Tom

For each 6 months Major Tom spends in the ISS, major Tom is 0.007 seconds younger than Bob.

## A realistic worldline: Langevin's travel



Image taken from Erik Gourgoulhon's book Special Relativity in General Frames

$$
\begin{array}{ll}
\text { for } t \in\left[0, \frac{T}{4}\right]: & x(t)=\frac{c T}{\alpha}\left[\sqrt{1+\alpha^{2}(t / T)^{2}}-1\right] \\
\text { for } t \in\left[\frac{T}{4}, \frac{3 T}{4}\right]: & x(t)=\frac{c T}{\alpha}\left[-\sqrt{1+\alpha^{2}(t / T-1 / 2)^{2}}+2 \sqrt{1+\frac{\alpha^{2}}{16}}-1\right]
\end{array}
$$

$$
\begin{equation*}
\text { for } t \in\left[\frac{3 T}{4}, T\right]: \quad x(t)=\frac{c T}{\alpha}\left[\sqrt{1+\alpha^{2}(t / T-1)^{2}}-1\right], \tag{2.20b}
\end{equation*}
$$

Eq. 2.20a rewritten

$$
\left(\alpha \frac{x}{c T}+1\right)^{2}-\left(\alpha \frac{t}{T}\right)^{2}=1
$$

2.20a-b-c: three hyperbolas cut and joint


## A realistic worldline: Langevin's travel



Image taken from Erik Gourgoulhon's book Special Relativity in General Frames

Compute worldline's tangent

$$
\mathrm{d} x=(-1)^{k} \frac{\alpha(t / T-k / 2)}{\sqrt{1+\alpha^{2}(t / T-k / 2)^{2}}} c \mathrm{~d} t
$$

Compute proper time

$$
c d \tau=\sqrt{-\left\|\dot{\boldsymbol{x}}^{2}\right\|_{M i n k}}
$$

$$
\mathrm{k}=1
$$

$$
c d \tau=\frac{d t}{\sqrt{1+\alpha^{2}(t / T-k / 2)^{2}}}
$$

Integrate this ODE to get

$$
\tau=\tau(t)
$$

and evaluate at initial and final events

$$
\frac{T^{\prime}}{T}=\frac{\tau(B)-\tau(A)}{t(B)-t(A)}=\frac{4}{\alpha} \operatorname{arcsinh}\left[\frac{\alpha}{4}\right] \leq 1
$$

## Mach's principle and rotation: Absolute or Relative?



Taken from The Forgotten Mystery of Inertia. American Scientiest.
Rotation is absolute!

Interestellar's docking scene.
Rotation is Relative!
Rotation is Relative!

## Mach's principle and Einstein's road to General Relativity

## Mach's principle:

inertial forces experienced by a body in nonuniform motion are determined by the quantity and distribution of matter in the universe.


## General Relativity

"Matter tells space how to bend, space tells matter how to move"


Geodesic equation
$\nabla_{\dot{\boldsymbol{x}}} \dot{\boldsymbol{x}}=0$
$\frac{d^{2} x^{\mu}}{d \tau^{2}}+\Gamma_{\alpha \beta}^{\mu} \frac{d x^{\alpha}}{d \tau} \frac{d x^{\beta}}{d \tau}=0$


